

Problem set #2 (Modern Physics)

04/16/2018 Provided by Masahito Oh-e

Solve the problems below. Describe the ways of thinking in English: only final solutions are not accepted. Make clear how you reach each solution.

Problem 1.

Calculate photon energies of the edges of the visible range spectrum at $\lambda=380$ nm, and $\lambda=770$ nm, respectively, with the units of J and eV.

Photon energy ϵ is expressed as $\epsilon=h\nu=\frac{hc}{\lambda}$, where h is the Planck's constant, c the speed of light, ν the frequency of light, and λ the wavelength of light. Therefore, for each wavelength $\lambda=380$ nm and $\lambda=770$ nm, the energies are calculated as

$$\epsilon_{380\text{nm}} = \frac{(6.626 \times 10^{-34}) \times (3.0 \times 10^8)}{380 \times 10^{-9}}$$

$$5.23105 \times 10^{-19}$$

$$\frac{5.23105 \times 10^{-19}}{(1.6 / 10^{-19})}$$

$$3.26941$$

$$\epsilon_{770\text{nm}} = \frac{(6.626 \times 10^{-34}) \times (3.0 \times 10^8)}{770 \times 10^{-9}}$$

$$2.58156 \times 10^{-19}$$

$$\frac{2.58156 \times 10^{-19}}{(1.6 / 10^{-19})}$$

$$1.61347$$

$$\epsilon(380 \text{ nm}) = 5.2 \times 10^{-19} \text{ J} = 3.3 \text{ eV}$$

$$\epsilon(770 \text{ nm}) = 2.6 \times 10^{-19} \text{ J} = 1.6 \text{ eV}$$

Problem 2.

Show that the Planck spectral distribution formula leads to the experimentally observed Stefan law for the total radiation emitted by a blackbody at all wavelengths. Stefan law: $E_{\text{total}} = aT^4 \text{ W} \cdot \text{m}^{-2} \cdot \text{K}^{-4}$, where T is the temperature. Use the relation $\int_0^{\infty} \frac{x^3}{(e^x-1)} dx = \frac{\pi^4}{15}$.

According to Planck's law, the total energy density is expressed as $u(\nu) = \frac{8\pi\nu^2}{c^3} \frac{h\nu}{\exp(h\nu/kT)-1}$, where h is the Planck's constant, c the speed of light, ν the frequency of light, T the temperature, and k the Boltzmann constant. Since Stefan's law is an expression for the total power per unit area radiated at

all frequencies, we must integrate the expression for $u(\nu)$ over all frequencies. If we make the change of variables $x=h\nu/kT$, we find $dx=h/kT(d\nu)$, and

$$u = \int_0^{\infty} u(\nu) d\nu = \frac{8\pi}{c^3} \frac{k^3 T^3}{h^2} \int_0^{\infty} \frac{(h\nu/kT)^3}{\exp(h\nu/kT)-1} d\nu = \frac{8\pi}{c^3} \frac{k^3 T^3}{h^2} \frac{kT}{h} \int_0^{\infty} \frac{x^3}{\exp(x)-1} dx = \frac{8\pi k^4 T^4}{c^3 h^3} \int_0^{\infty} \frac{x^3}{\exp(x)-1} dx = \frac{8\pi^4 k^4}{15 c^3 h^3} T^4$$

Problem 3.

Suppose that light of total intensity $1.0 \mu\text{W}/\text{cm}^2$ falls on a clean iron sample 1.0 cm^2 in area.

Assume that the iron sample reflects 96% of the light and that only 3.0% of the absorbed energy lies in the violet region of the spectrum above the threshold frequency.

- What intensity is actually available for the photoelectric effect?
- Assuming that all the photons in the violet region have an effective wavelength of 250 nm, how many electrons will be emitted per second?
- Calculate the current in the photo-tube in amperes.
- If the cutoff frequency is $f_0 = 1.1 \times 10^{15}$ Hz, find the work function, ϕ , for iron.
- Find the stopping voltage for iron if photo-electrons are produced by light with $\lambda=250$ nm.

(a) Because only 4.0% of the incident energy is absorbed, and only 3.0% of this energy is able to produce photo-electrons, the intensity available is

$$I = (0.030) \times (0.040) I_0 = (0.030) \times (0.040) \times (1.0 \mu\text{W}/\text{cm}^2) = 1.2 \text{ nW}/\text{cm}^2.$$

(b) For an efficiency of 100%, one photon of energy, $h\nu$, will produce one electron, so

of electrons /s =

$$\frac{1.2 \times 10^{-9} \text{ W}}{h\nu} = \frac{\lambda \times 1.2 \times 10^{-9}}{hc} = \frac{250 \times 10^{-9} \times 1.2 \times 10^{-9}}{6.626 \times 10^{-34} \times 3.0 \times 10^8}$$

$$1.50921 \times 10^9$$

of electrons / s = 1.5×10^9

(c) The current i can be calculated by $i=e \times (\# \text{ of electrons/s})$

$$i = (1.602 \times 10^{-19} \text{ C}) \times (1.5 \times 10^9 \text{ electrons/s}) = 2.4 \times 10^{-10} \text{ A}$$

(d) Work function ϕ can be expressed as hf_0 . Therefore,

$$\phi = hf_0 = (4.14 \times 10^{15} \text{ eV} \cdot \text{s}) (1.1 \times 10^{15} \text{ s}^{-1}) = 4.5 \text{ eV}$$

(e) From the photoelectric equation, the energy of photo-electrons can be calculated by

$$e V_s = hf - \phi = \frac{hc}{\lambda} - \phi = \frac{(4.14 \times 10^{15} \text{ eV} \cdot \text{s}) \times (3.0 \times 10^8 \text{ m/s})}{250 \times 10^{-9} \text{ m}} - 4.5 \text{ eV} = 0.46 \text{ eV}$$

Problem 4.

X-rays of wavelength $\lambda=0.200$ nm are aimed at a block of carbon. The scattered x-rays are observed at an angle of 45.0° to the incident beam. Calculate the increased wavelength of the scattered x-rays at this angle.

The change in wavelength expected by a photon that is scattered through the angle θ by a particle of rest mass m is expressed by

$$\Delta\lambda = \lambda - \lambda_0 = \frac{h}{mc} (1 - \cos\theta). \text{ Taking } \theta=45^\circ, \text{ we find}$$

$$\Delta\lambda = \frac{6.63 \times 10^{-34} \text{ J}\cdot\text{s}}{(9.11 \times 10^{-31} \text{ kg})(3.00 \times 10^8 \text{ m/s})} (1 - \cos 45.0^\circ) = 7.11 \times 10^{-13} \text{ m} = 0.00071 \text{ nm}$$

Hence, the wavelength of the scattered x-ray at this angle is

$$\lambda = \Delta\lambda + \lambda_0 = 0.200711 \text{ nm}$$

Problem 5.

An electron of charge q and mass m is accelerated from rest through a small potential difference $V=50 \text{ V}$. Calculate de Broglie wavelength λ , assuming that the particle is nonrelativistic.

When a charge is accelerated from rest through a potential difference V , its gain in kinetic energy, $\frac{1}{2} mv^2$, must equal the loss in potential energy qV . That is, $\frac{1}{2} mv^2 = qV$. Because $p = mv$, we can express this in the form

$$\frac{p^2}{2m} = qV \text{ or } p = \sqrt{2mqV}$$

Substituting this expression for p into the de Broglie relation $\lambda = h/p$ gives

$$\lambda = (h/p) = \frac{h}{\sqrt{2mqV}}. \text{ Therefore, the de Broglie wavelength of an electron accelerated through } 50 \text{ V is}$$

$$\lambda = h/\text{Sqrt}[2mqV] = \frac{6.63 \times 10^{-34} \text{ J}\cdot\text{s}}{\sqrt{2(9.11 \times 10^{-31} \text{ kg})(1.6 \times 10^{-19} \text{ C})(50 \text{ V})}} = 1.7 \times 10^{-10} \text{ m} = 1.7 \text{ \AA}$$

Problem 6.

Consider a $2 \mu\text{g}$ mass traveling with a speed of 10 cm/s . If the particle's speed is uncertain by 1.5%, what is its uncertainty in position? Use the uncertainty relation: $\Delta x \Delta p_x \geq h$.

With $m=2 \times 10^{-6} \text{ g} = 2 \times 10^{-9} \text{ kg}$ and $v = 10 \text{ cm/s} = 0.1 \text{ m/s}$, the uncertainty in speed is $\Delta v = (0.015)(0.1 \text{ m/s}) = 1.5 \times 10^{-3} \text{ m/s}$.

Thus, the uncertainty in position is given by

$$\Delta x = \frac{h}{\Delta p} = \frac{h}{m\Delta v} = \frac{6.63 \times 10^{-34} \text{ J}\cdot\text{s}}{(2 \times 10^{-9} \text{ kg})(1.5 \times 10^{-3} \text{ m/s})} = 2.21 \times 10^{-22} \text{ m}.$$